

Time Delay Interferometry using the UF LISA Benchtop Simulator

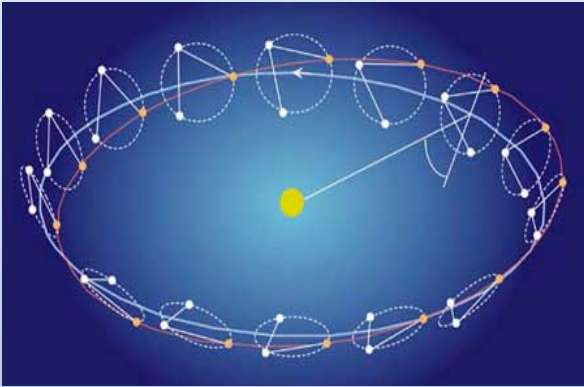
Rachel J. Cruz



6th International LISA Symposium

June 21, 2006

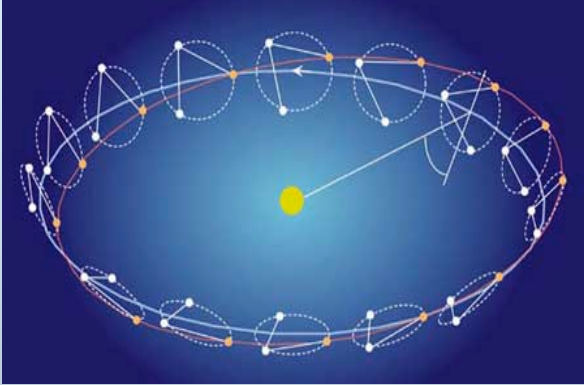
Laser Phase Noise




LISA needs to be able to see length changes of ~ 10 pm corresponding to a required frequency stability of less than $6 \times 10^{-7} \text{ Hz}/\sqrt{\text{Hz}}$.

- Pre-stabilization
 - Reference cavity
- Arm-locking
 - Use the LISA arm as a stable reference
- Time Delay Interferometry (TDI)

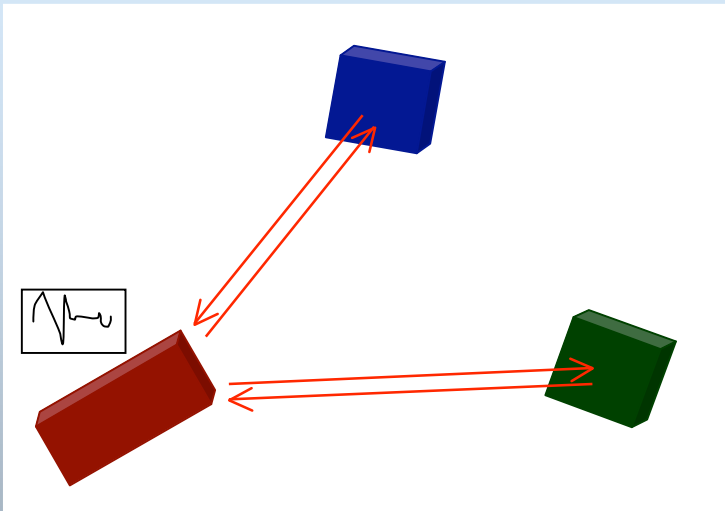
Laser Phase Noise



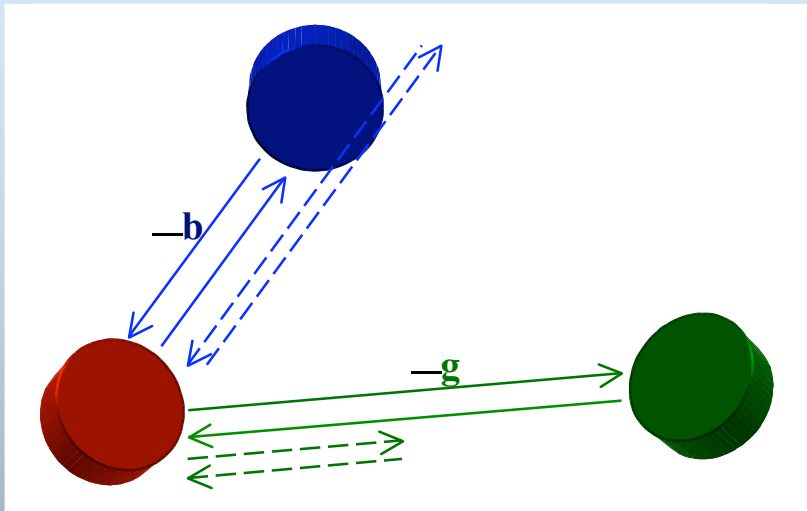
LISA needs to be able to see length changes of ~ 10 pm corresponding to a required frequency stability of less than $6 \times 10^{-7} \text{ Hz}/\sqrt{\text{Hz}}$.

- Pre-stabilization
 - Reference cavity
- Arm-locking  Ira Thorpe- LISA System & Technology Session Friday morning
 - Use the LISA arm as a stable reference
- Time Delay Interferometry (TDI)

Time Delay Interferometry



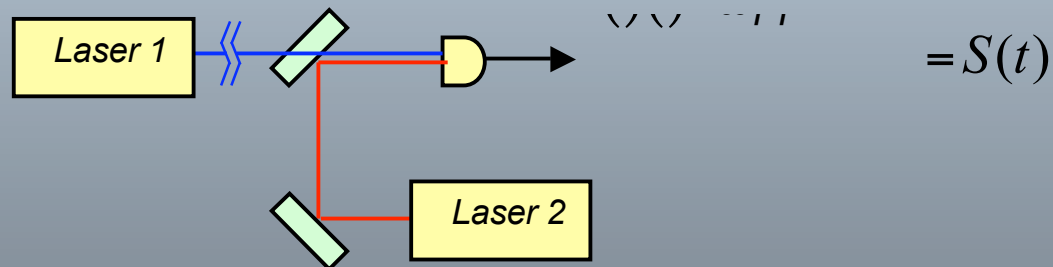
Time Delay Interferometry



First Generation X-combination:

$$S_b(t) - S_g(t) - S_b(t - 2_{-g}) + S_g(t - 2_{-b})$$

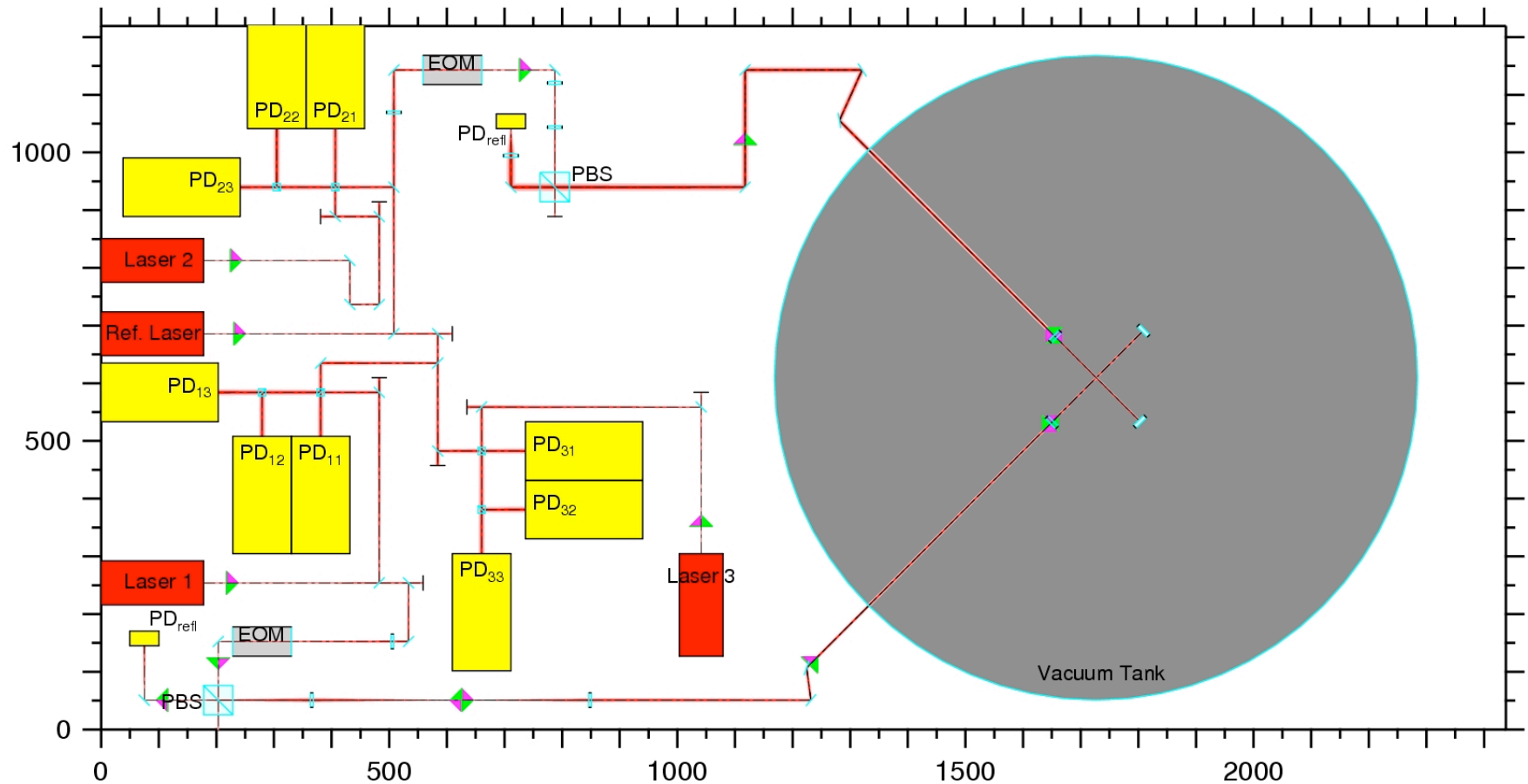
cancels laser phase noise



Additional terms can be added to compensate for transponder noise, bench noise, and spacecraft motion

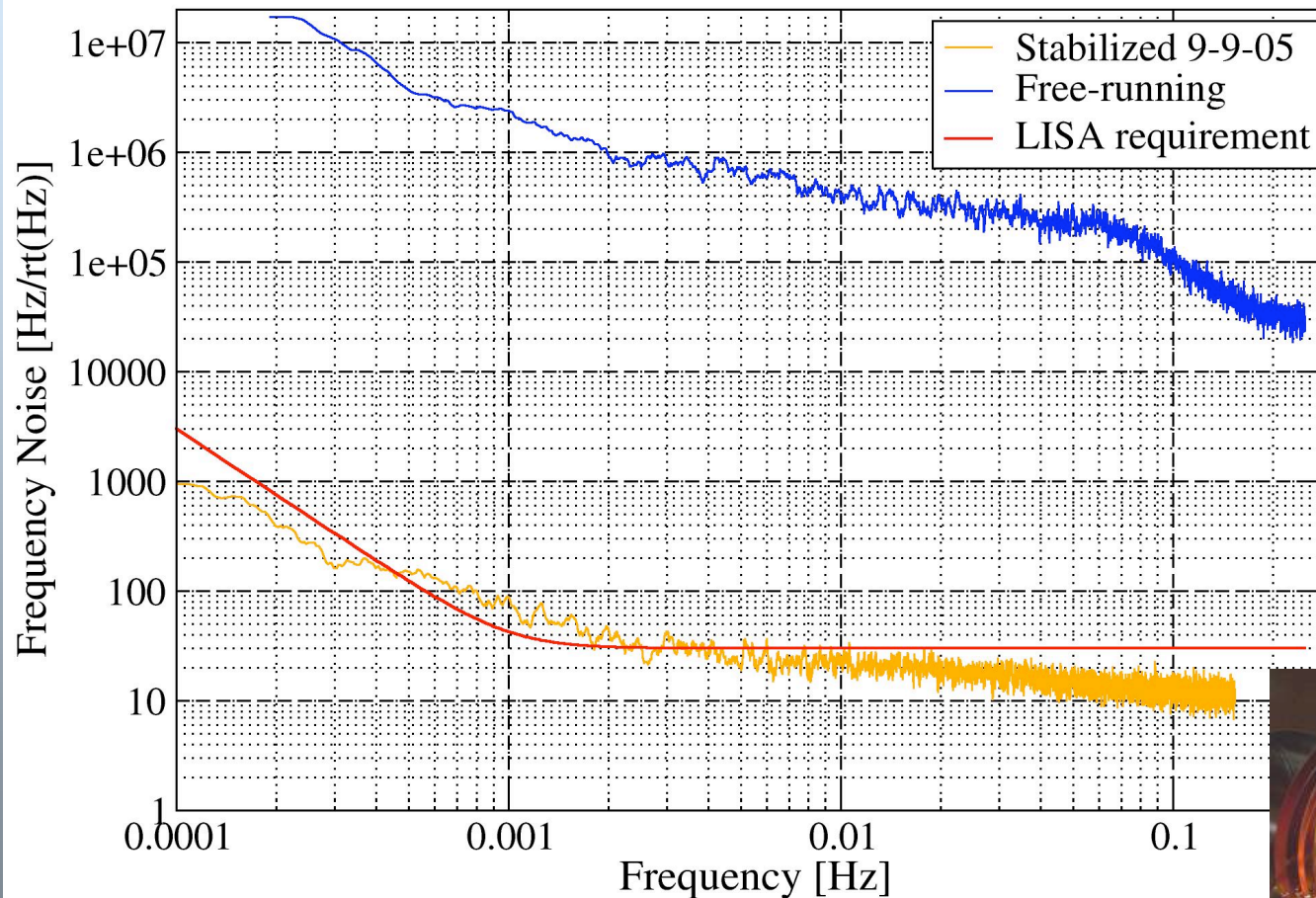
UF LISA Banchtop Simulator

Our goal is to create a facility to investigate experimentally aspects of LISA. Currently, we are looking at TDI, arm-locking, laser communication, and material and bond stability studies.

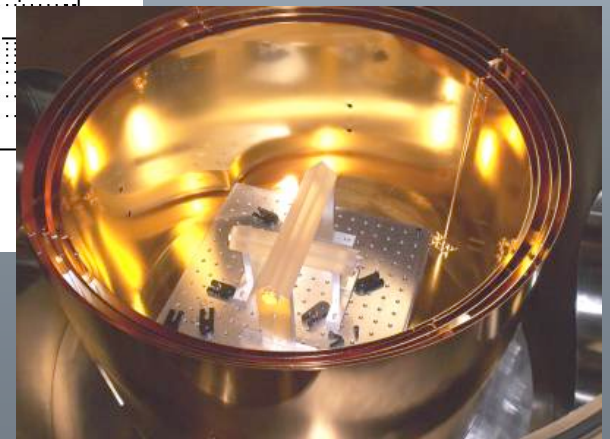


Laser frequency stabilization

Free-running vs. Cavity Stabilized Laser



Two Zerodur
spacers with
optically-
contacted
mirrors

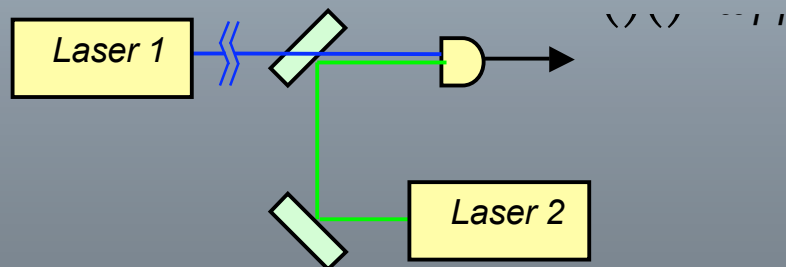
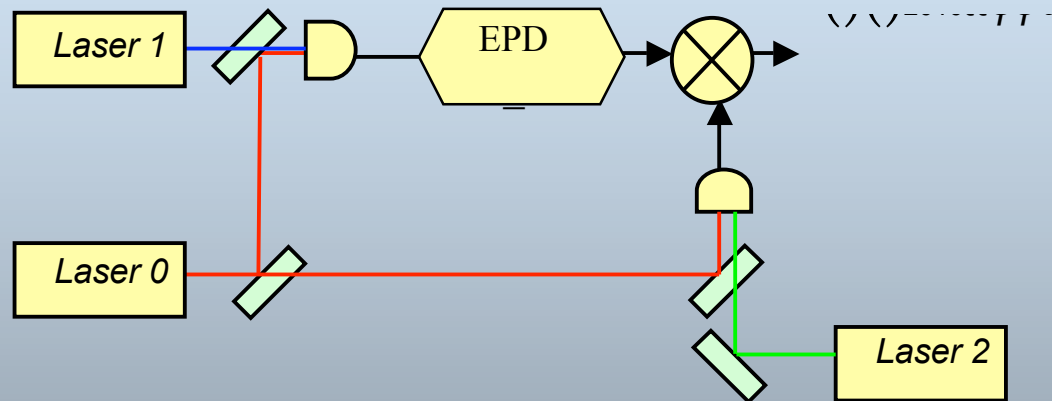


Electronic Phase Delay

We delay the optical signals electronically to create the long LISA arm in the lab.

EPD

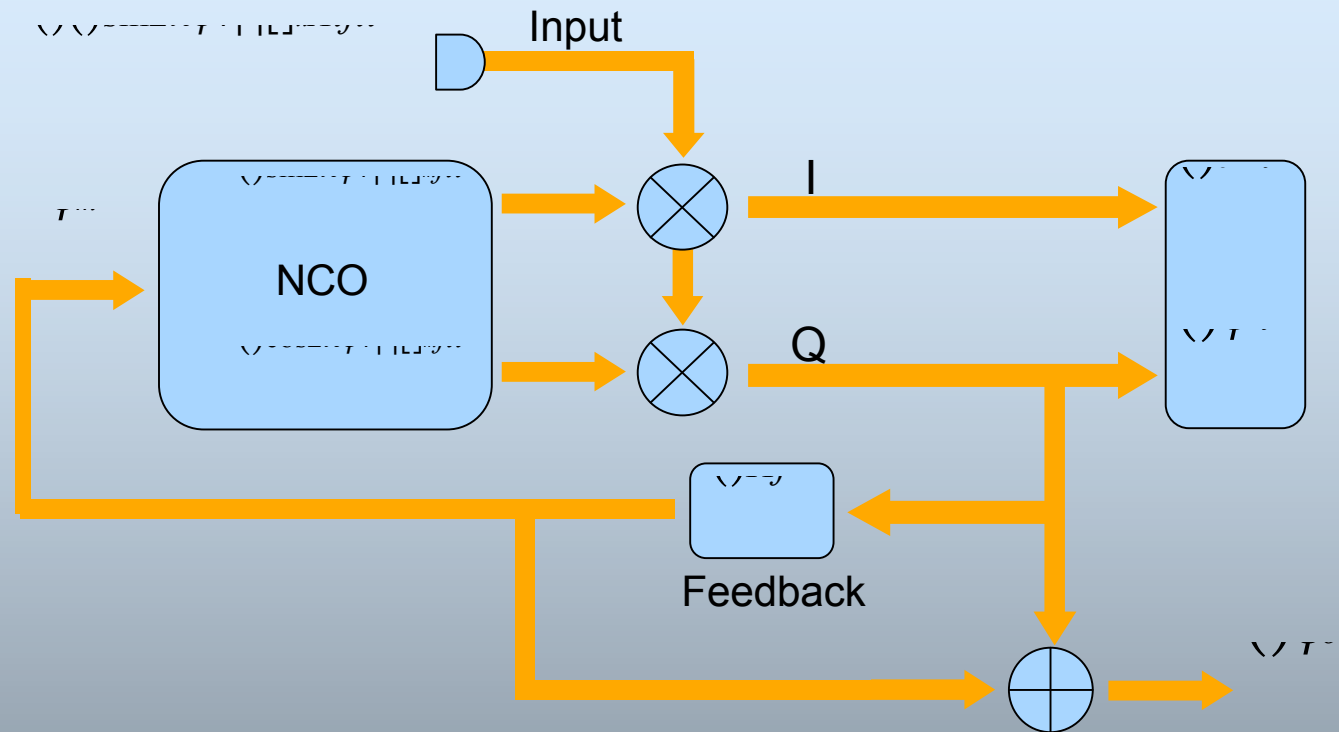
- Beat note is electronically delayed
- Reference laser is required



LISA

- LISA arm delays phase of light
- Beat note formed with 2nd laser

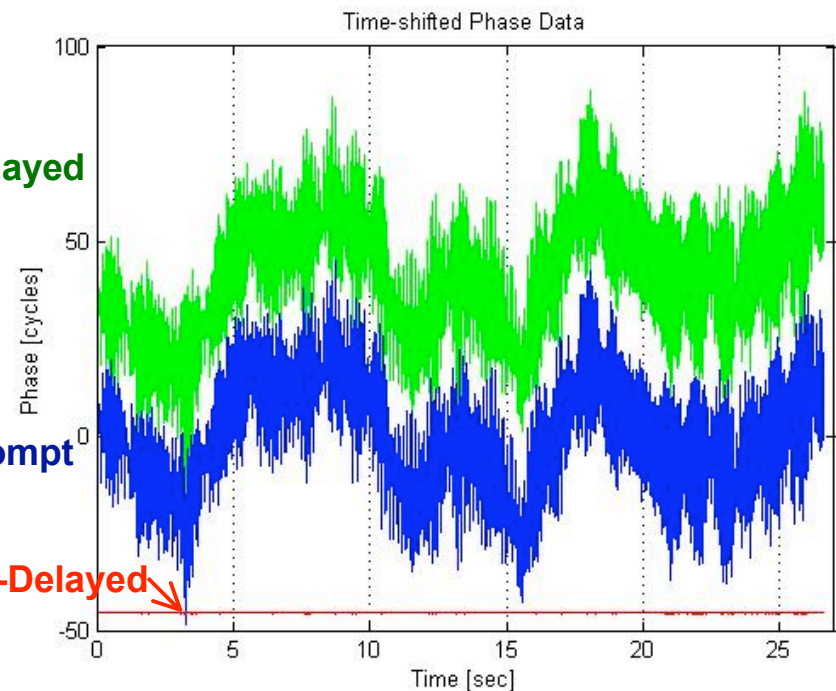
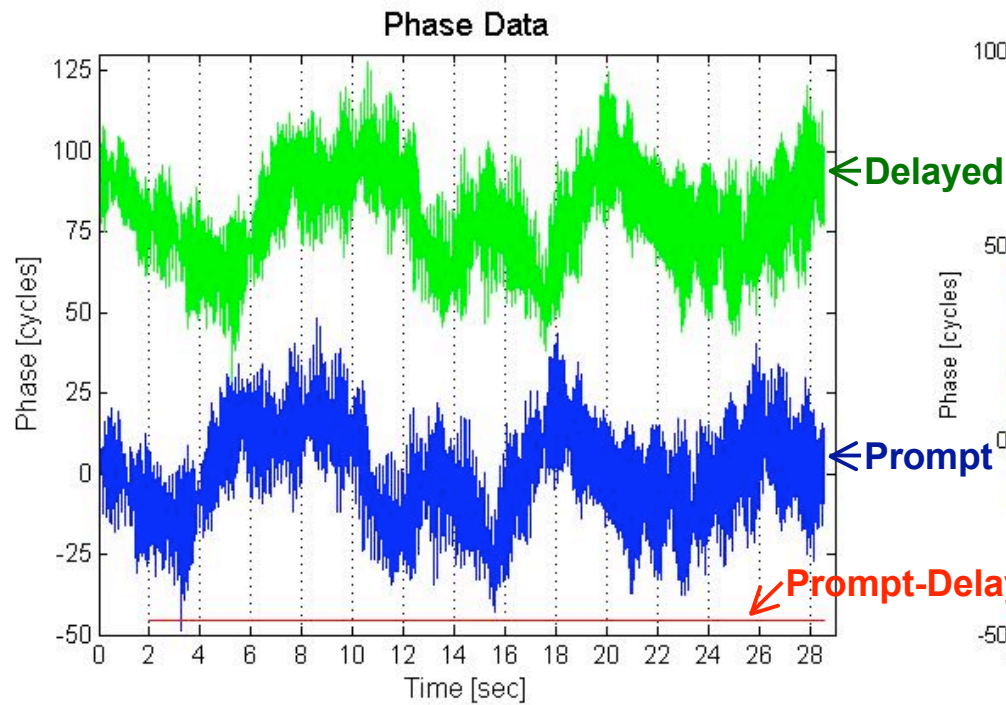
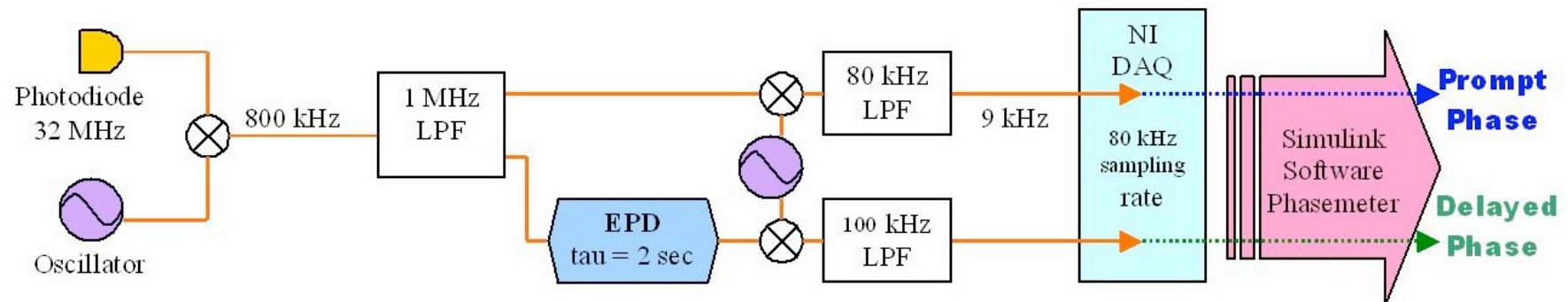
LISA-like Post-process Phasemeter



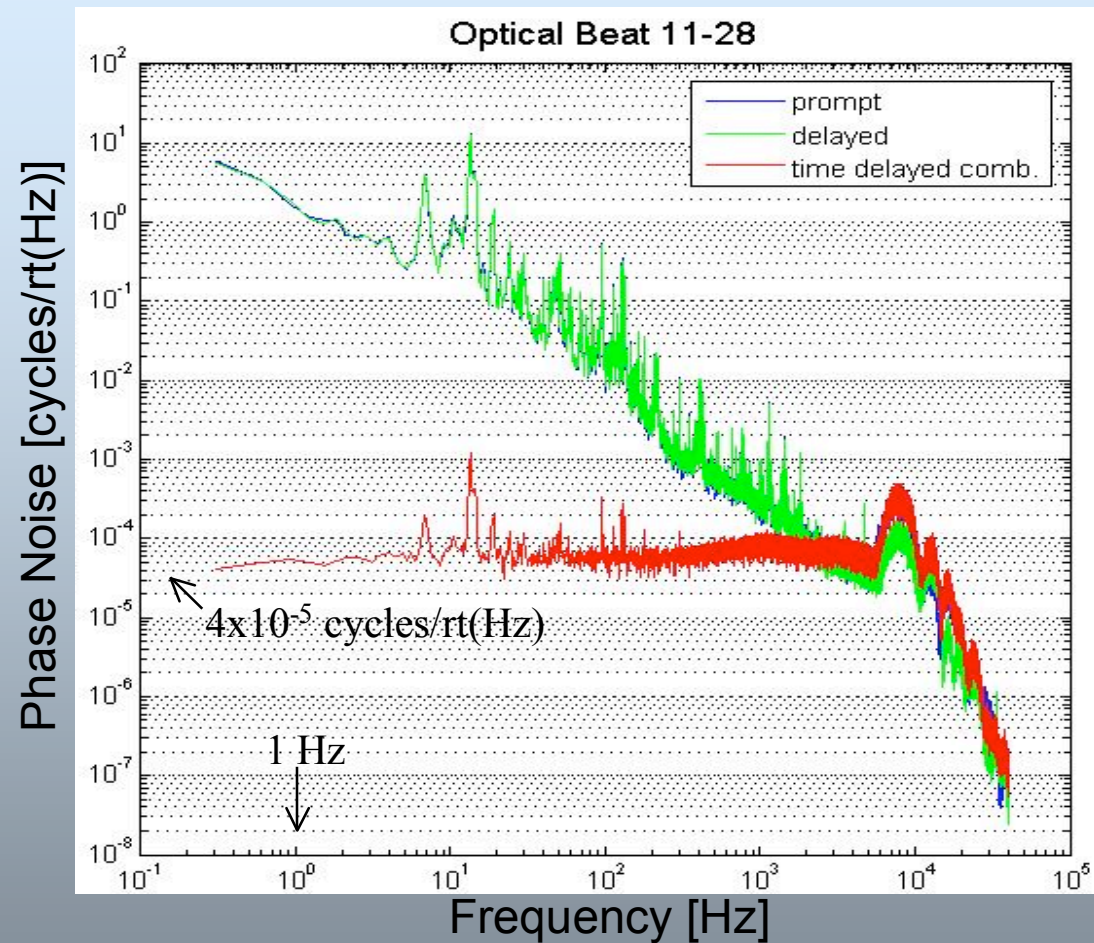
I used a post-process/software version in Simulink

- requires 10 kHz signals \diamond demodulation stage in experiment
- requires sampling at 80 kHz \diamond limits data runs to roughly a minute

Initial Optical Experiment



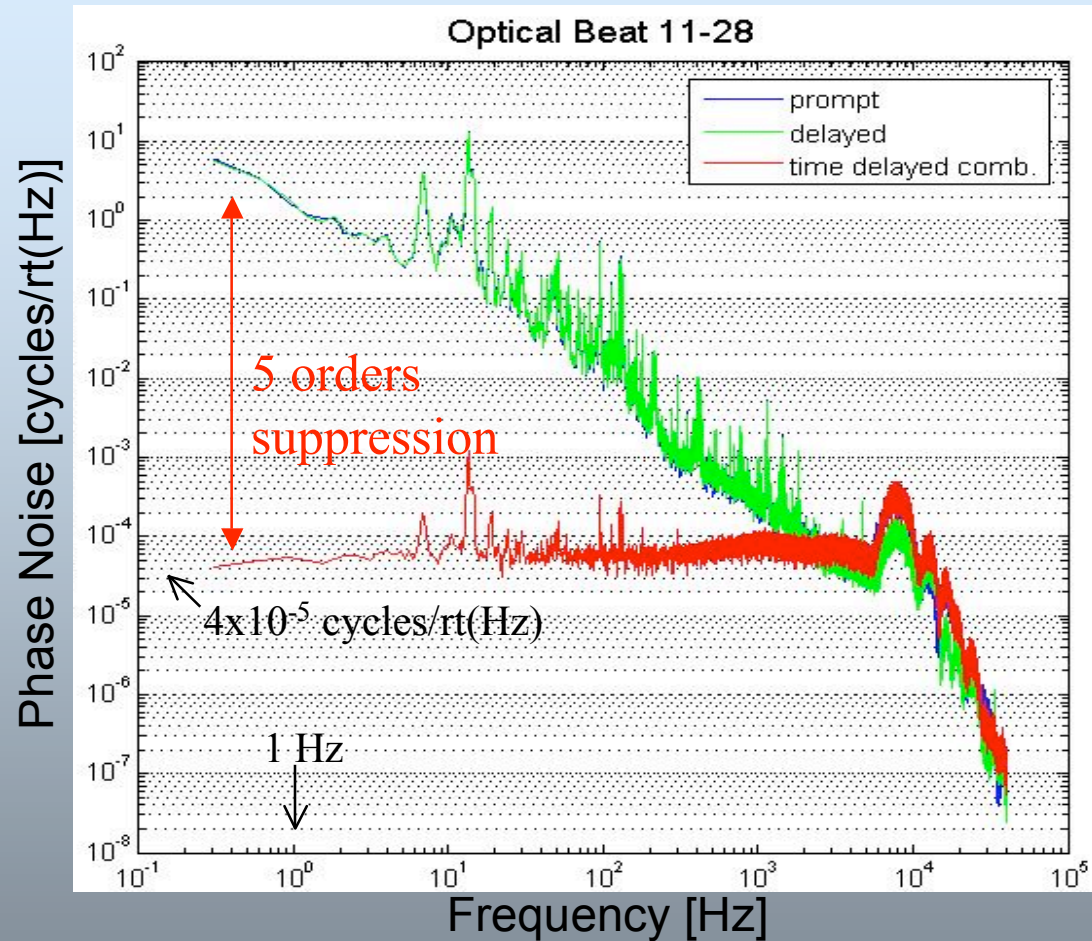
Initial Optical Experiment Results



Frequency range currently limited by sampling of post-process phasemeter.

We are less than an order of magnitude from LISA requirements!

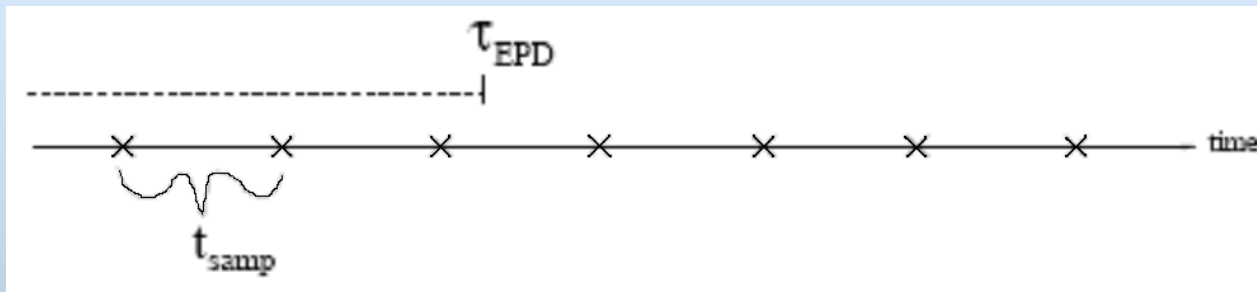
Initial Optical Experiment Results



Frequency range currently limited by sampling of post-process phasemeter.

We are less than an order of magnitude from LISA requirements!

Timing Error



The delay time of the EPD, just as the optical delay time of the LISA arm, will not fall exactly at one of the sampling points of the data stream.

Define the timing error as: $\Delta\tau = |\tau_{EPD} - \tau_{\text{shift}}|$

$$\Delta\tau_{\text{max}} = \frac{1}{2}t_{\text{samp}}$$

Suppression limits due to timing error

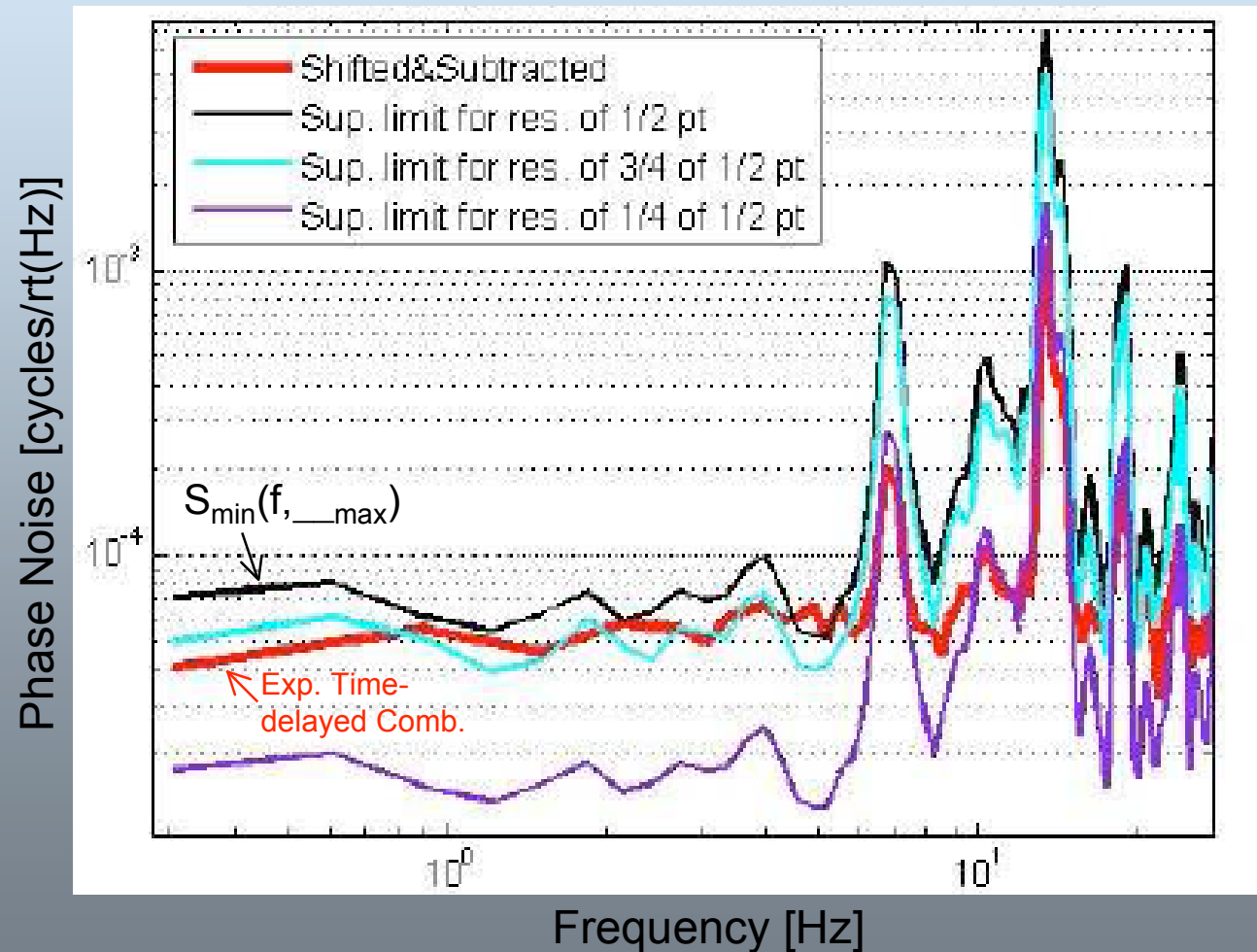
$$S(t) = \phi_p(t) - \phi_p(t - \Delta\tau)$$

$$|\tilde{S}_{\min}(f, \Delta\tau)| = 2\sin(\pi f \Delta\tau) |\tilde{\phi}_p(f)|$$

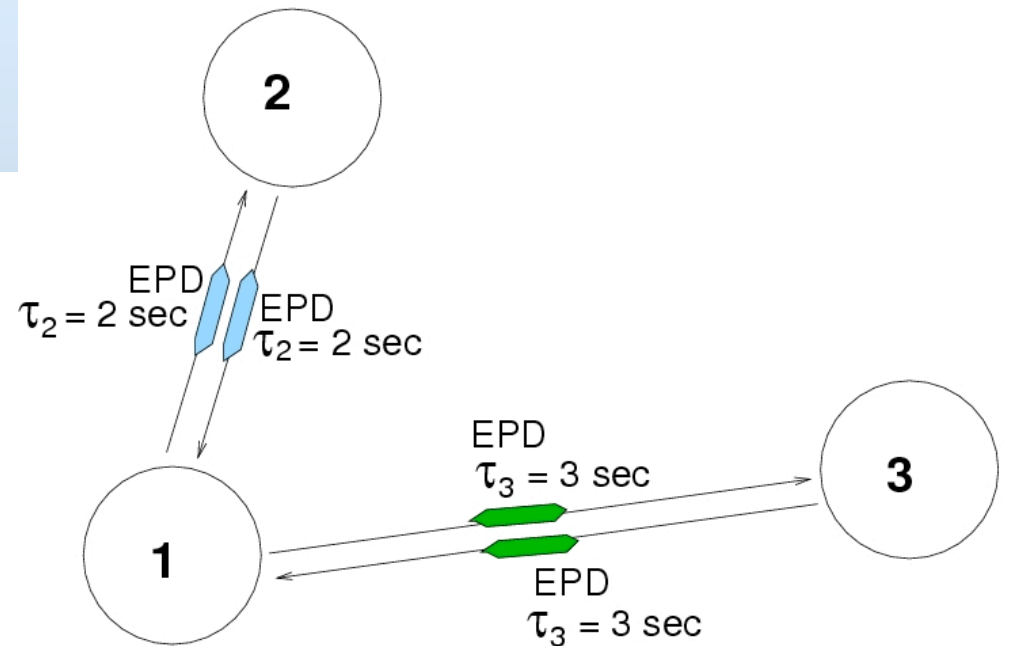
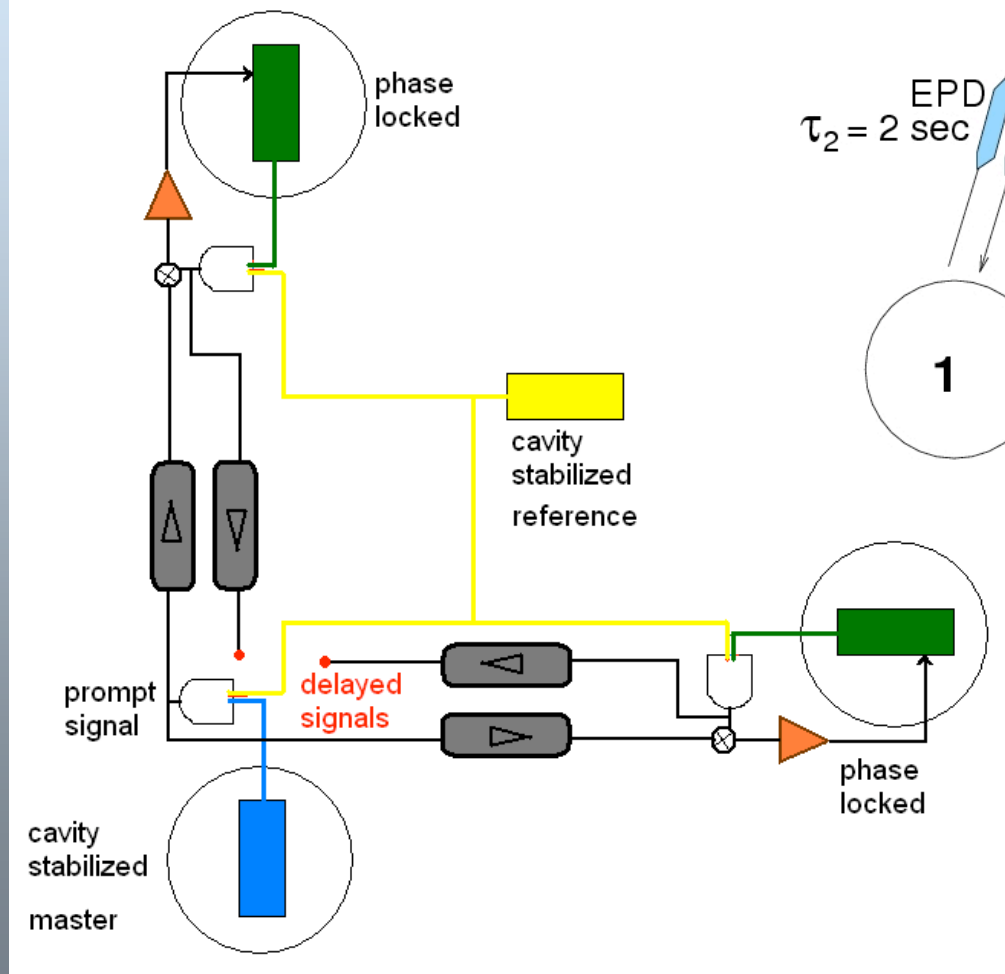
- The timing error in the experiment $< \text{__}_{\max} = \text{__}_{t_{\text{samp}}} = 6.25 \text{ _sec}$

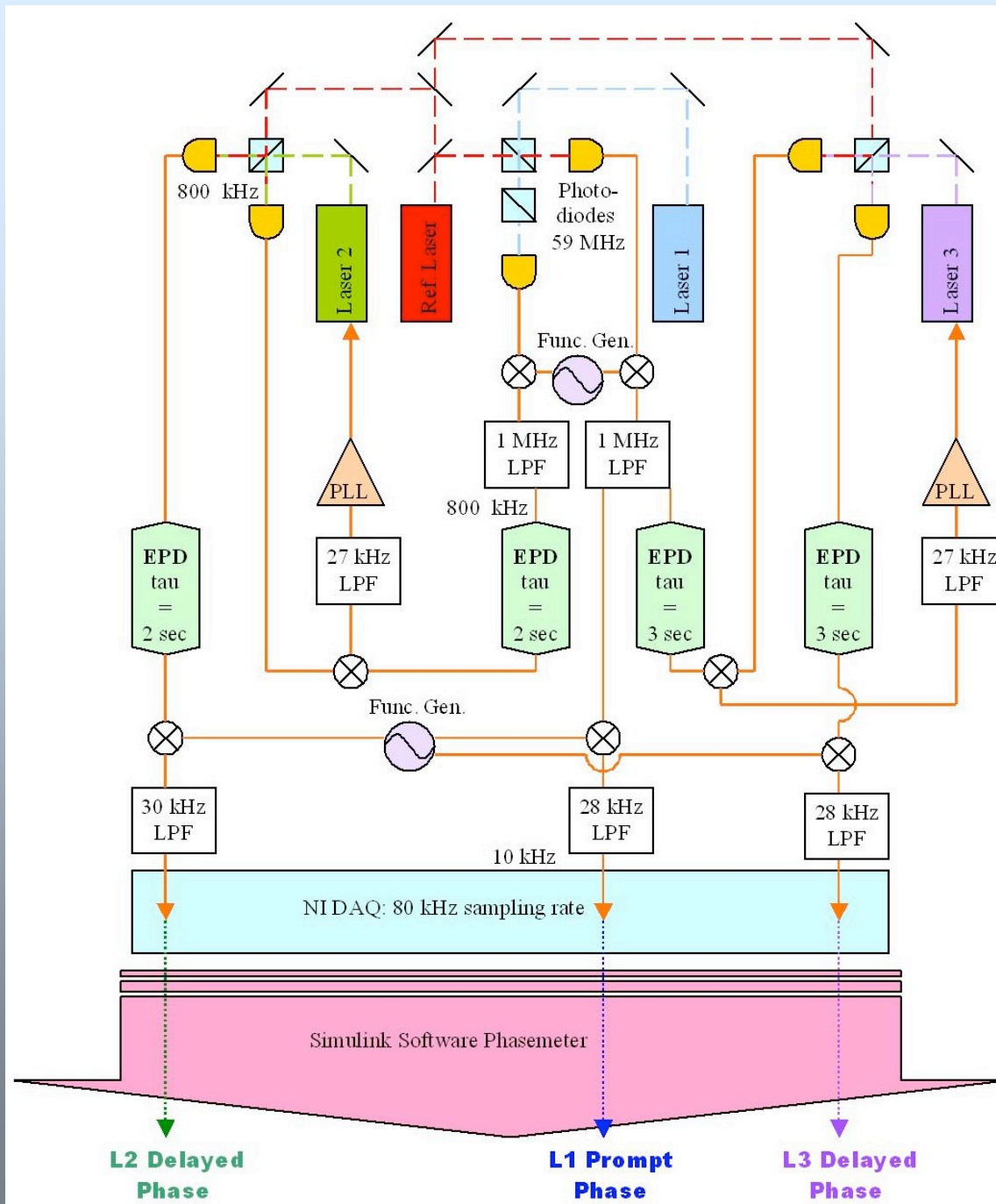
- Interpolation can be used to reduce the timing error

- Experimental results appear to hit another noise source at $\sim 5 \times 10^{-5}$ cycles/rt(Hz)



Two-Arm Experiment

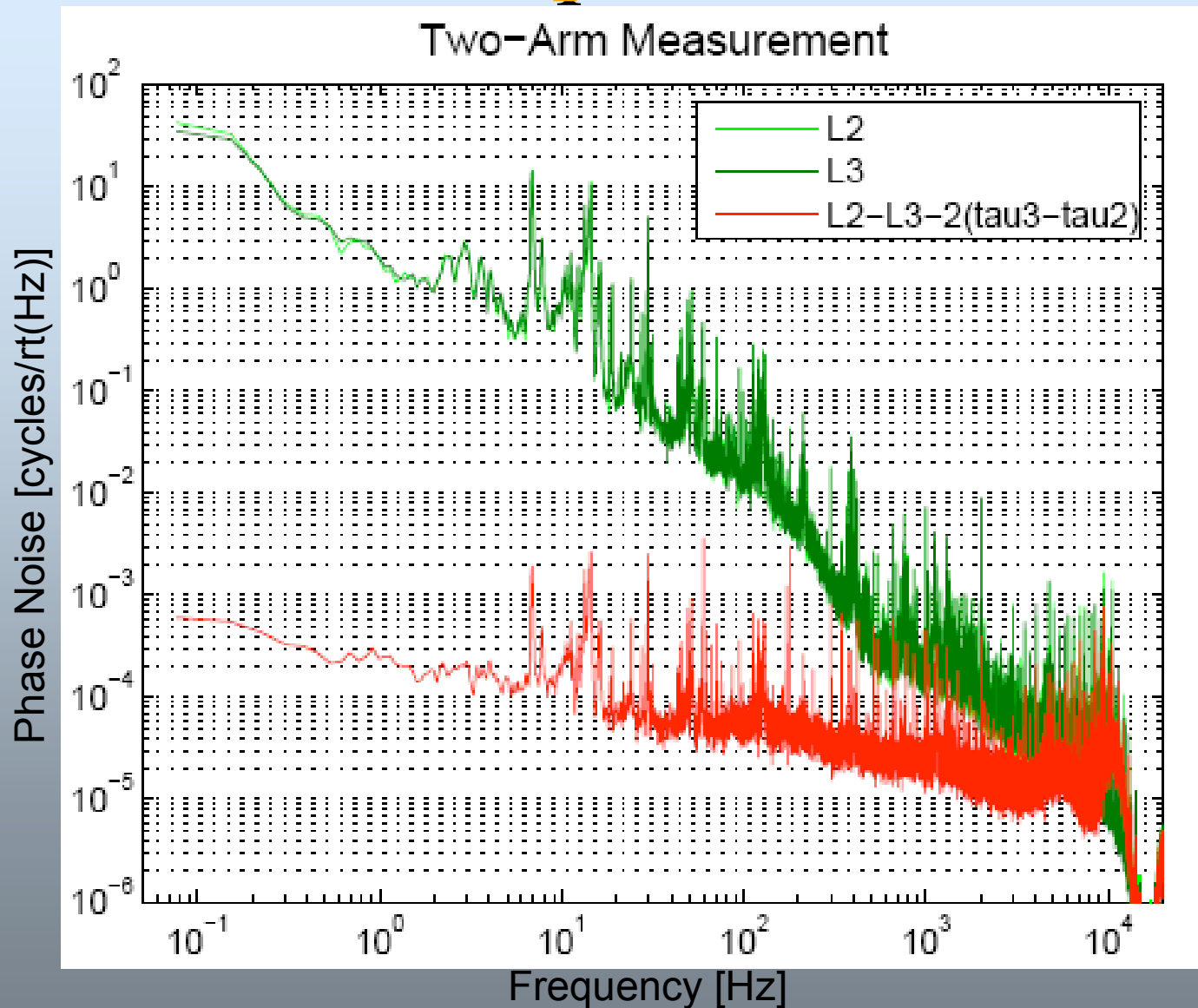




Benchtop Simulator during Two-Arm Experiment

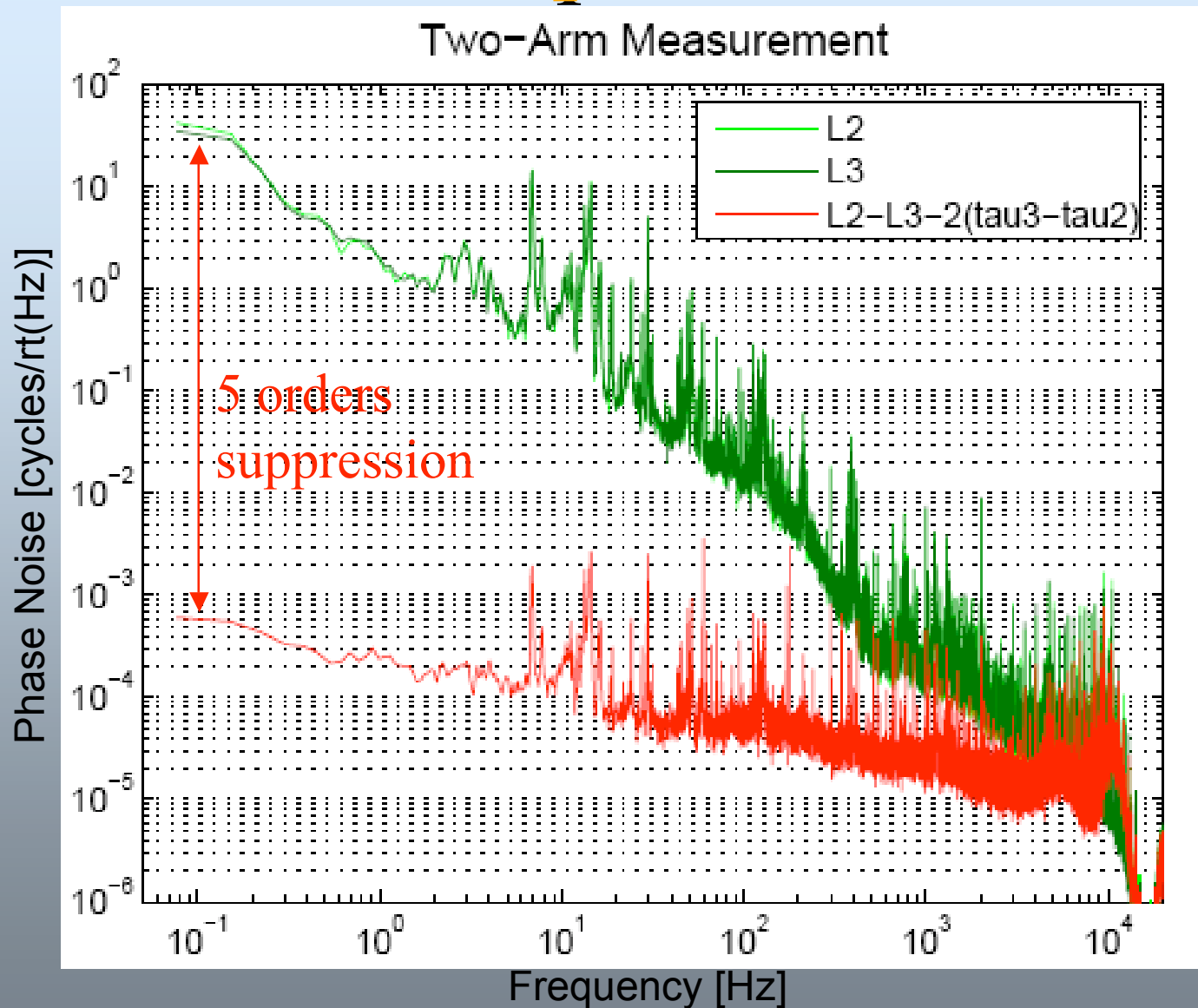


Two-Arm Experiment Results



Results currently limited by PLL performance

Two-Arm Experiment Results



Results currently limited by PLL performance



Summary & Outlook

- One Signal experiment:
 - Less than one order of magnitude above LISA requirement @ 100 mHz
- Two-arm TDI experiment:
 - 5 orders of magnitude noise suppression @100 mHz
 - Less than 2 orders of mag above LISA requirements
 - Limited by Phase lock loops
- Real Time phasemeter under development
 - Allows long data runs to get into the LISA band
 - Develop interpolation techniques
 - Enables Noise hunting
- Build-up full LISA simulator
 - With GW, Doppler shifts, Clock noise, split OPL...

www.phys.ufl.edu/research/lisa/

The UF LISA Team:

Guido Mueller, David Tanner, Sergei Klimenko,
Ira Thorpe, Michael Hartman, Rodrigo Delgadillo,
Shawn Mitryk, Aaron Worley, Gabriel Boothe,
Sridhar Reddy, Alix Preston, and Yinan Yu

Thanks also to Wan Wu (UF-LIGO).

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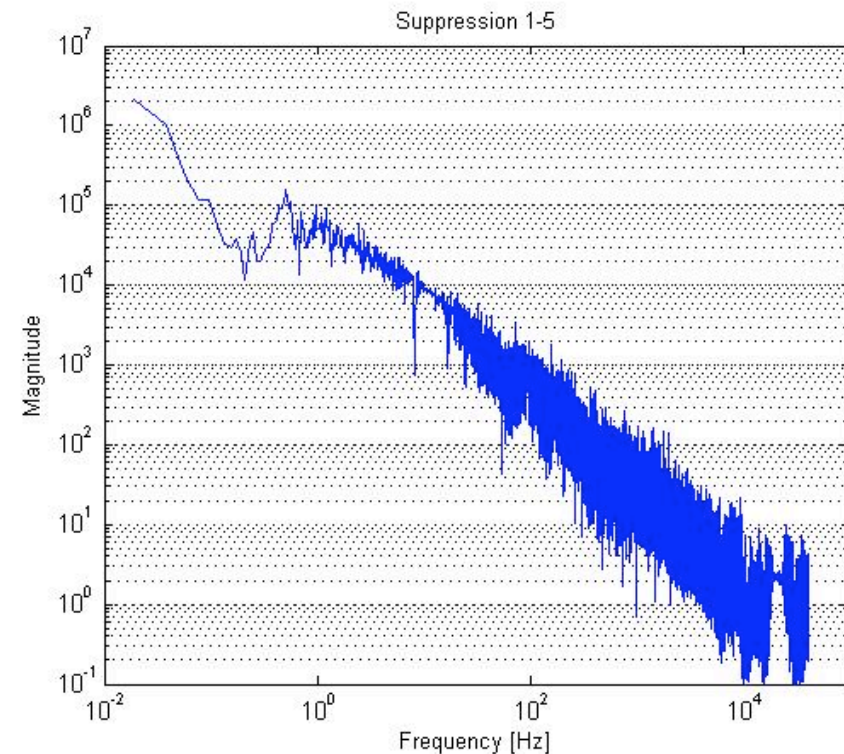
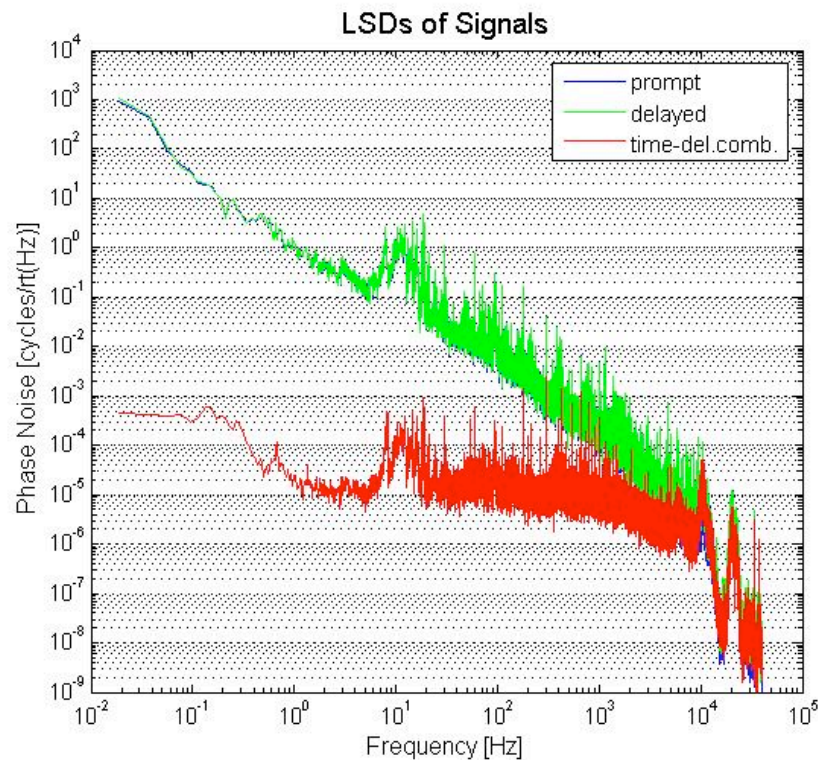
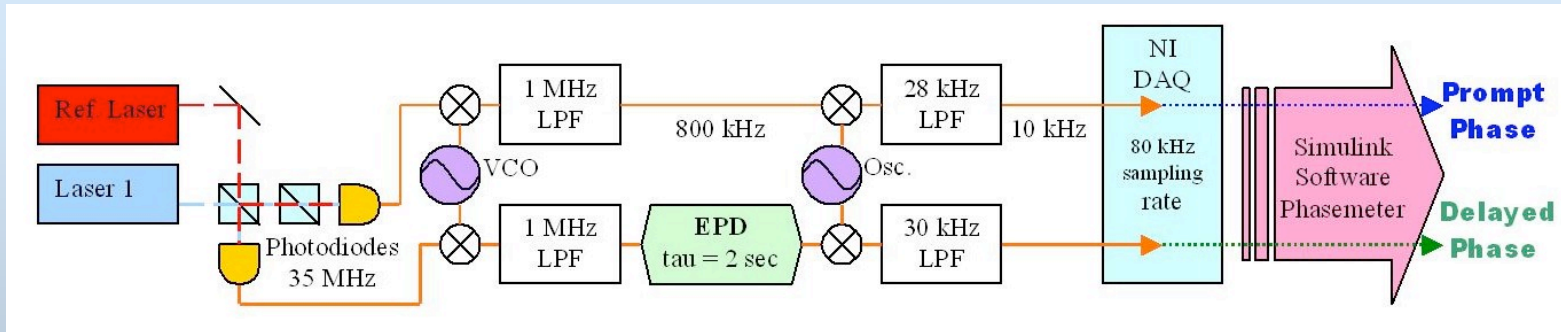
NASA Harriett G. Jenkins

Predoctoral Fellowship Program

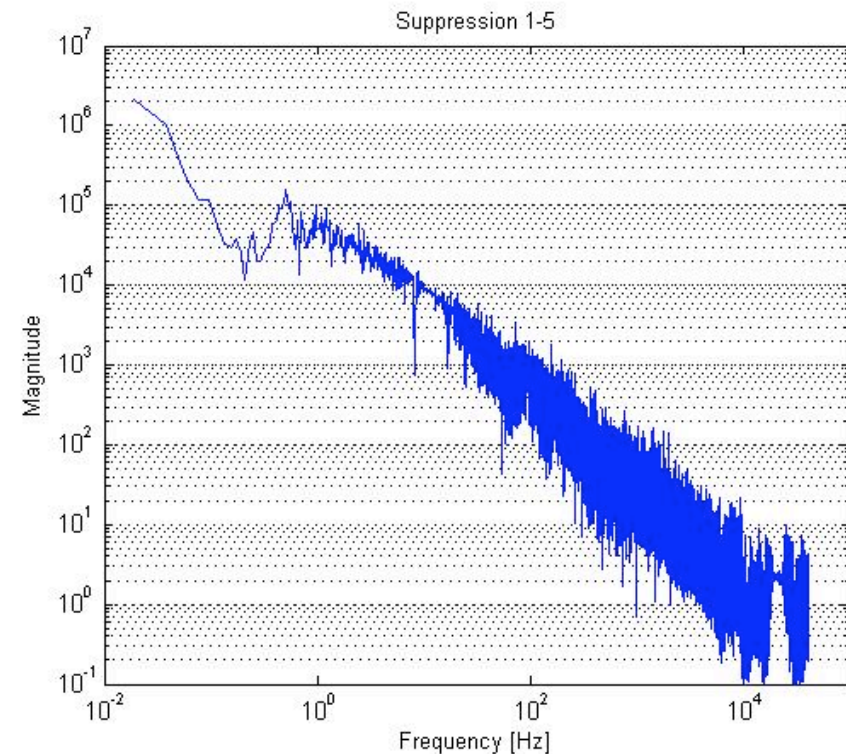
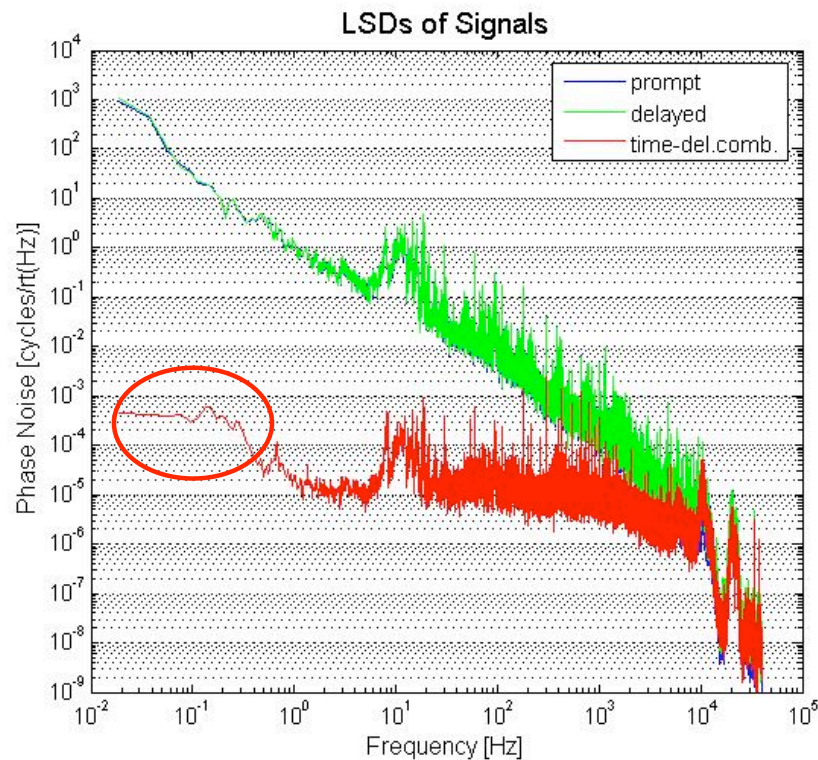
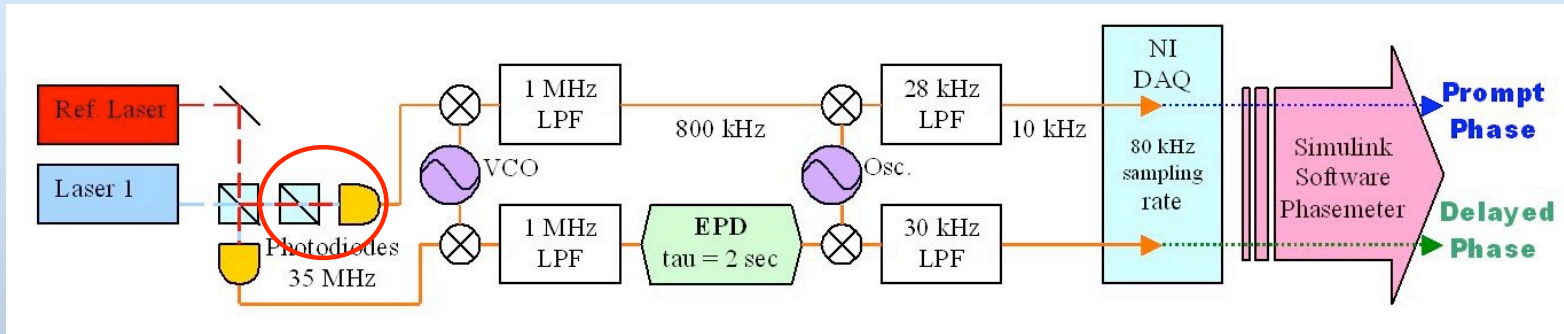
NASA



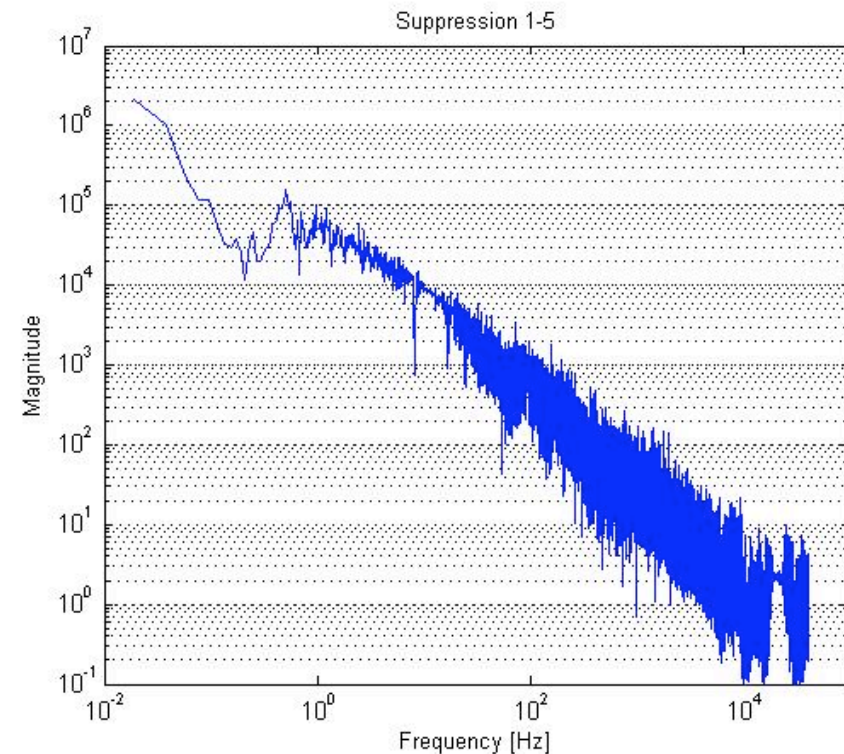
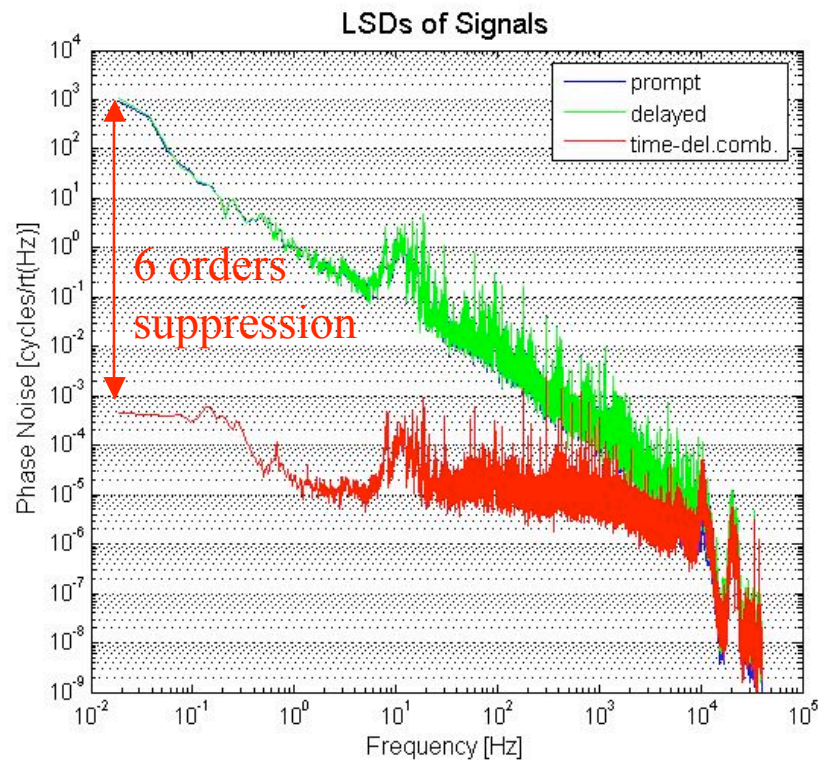
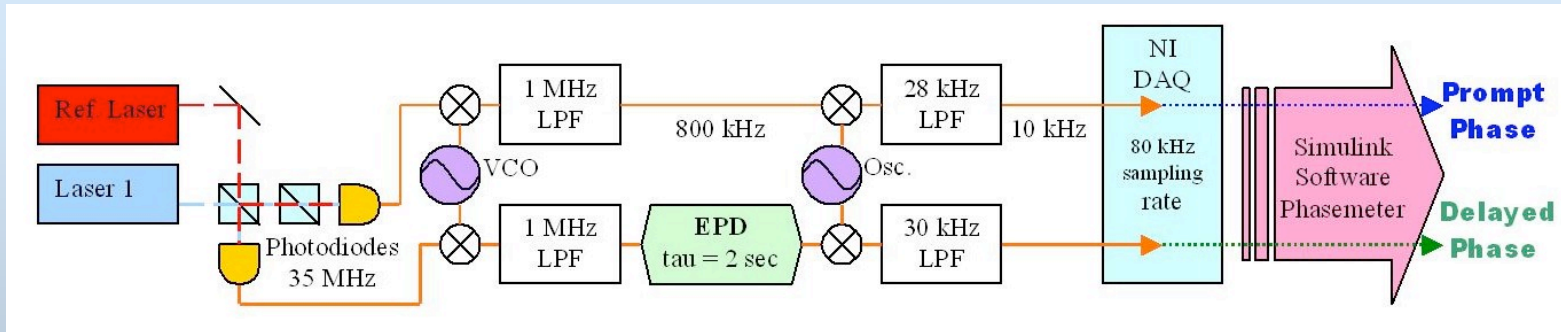
Optically-split Experiment



Optically-split Experiment



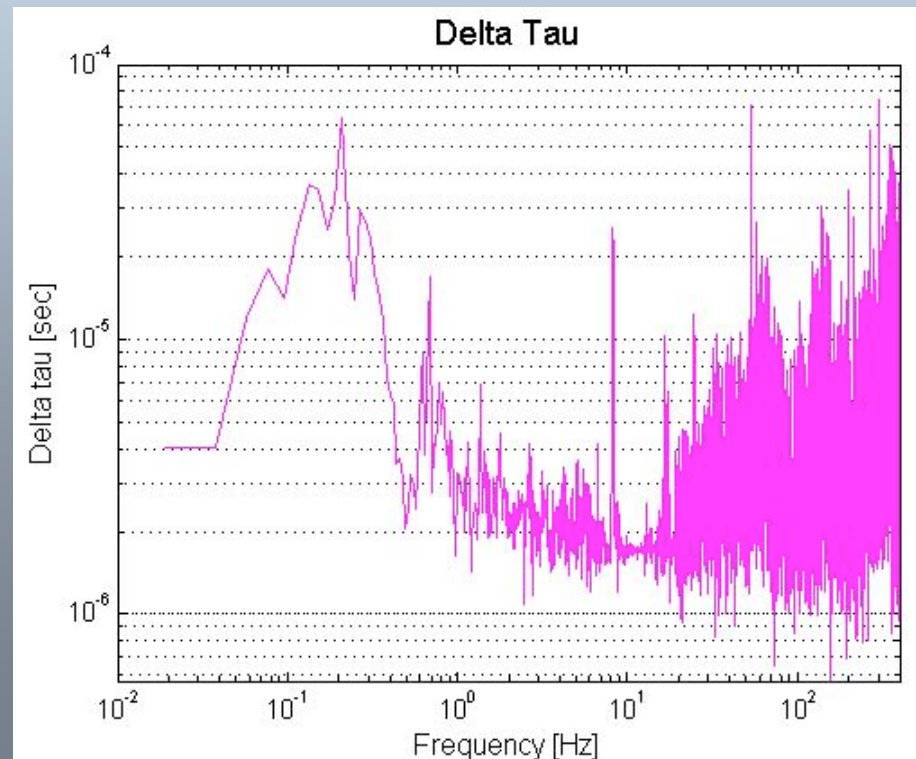
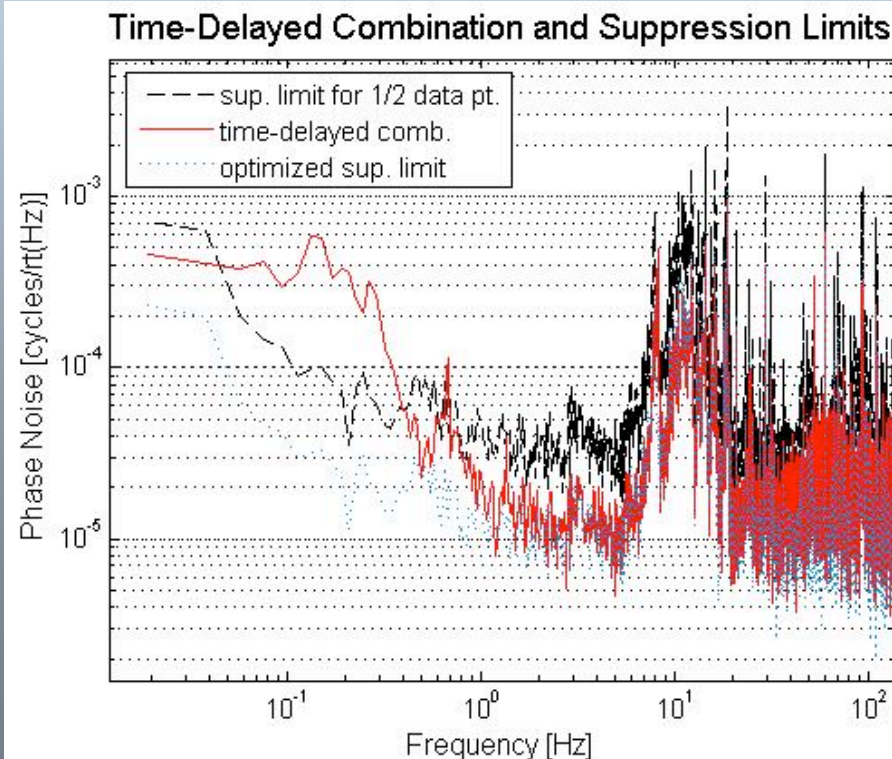
Optically-split Experiment



Timing Resolution Suppression Limit & Timing Error

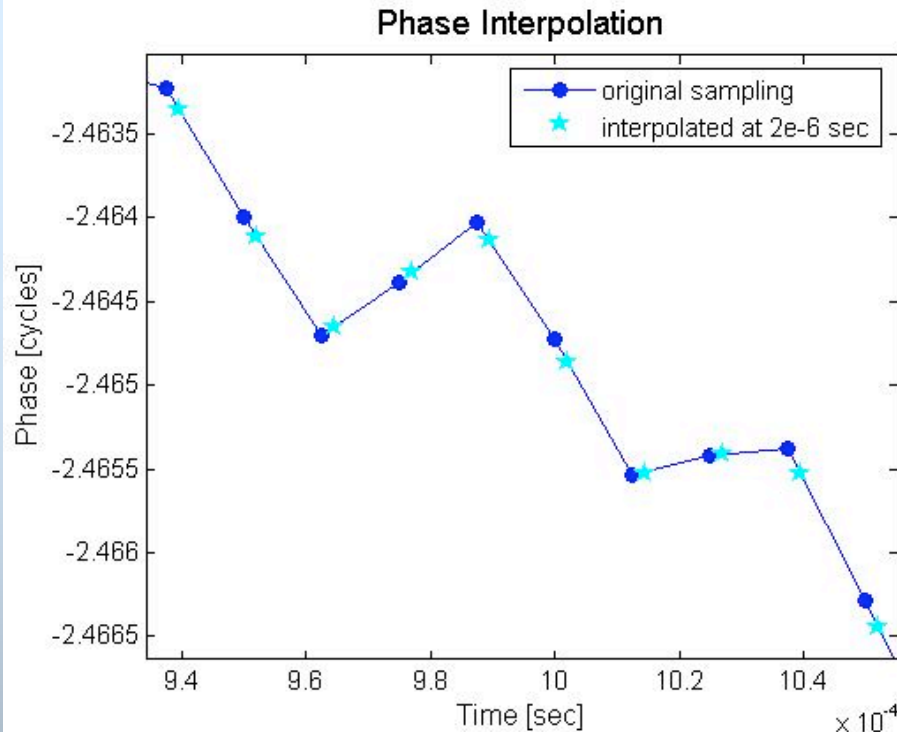
$$|\tilde{S}(f)| = 2\sin(\pi f \Delta\tau) |\tilde{\phi}_p(f)|$$

$$\Delta\tau = \frac{1}{\pi f} \sin^{-1} \left[\frac{|\tilde{S}(f)|}{2|\tilde{\phi}_p(f)|} \right]$$

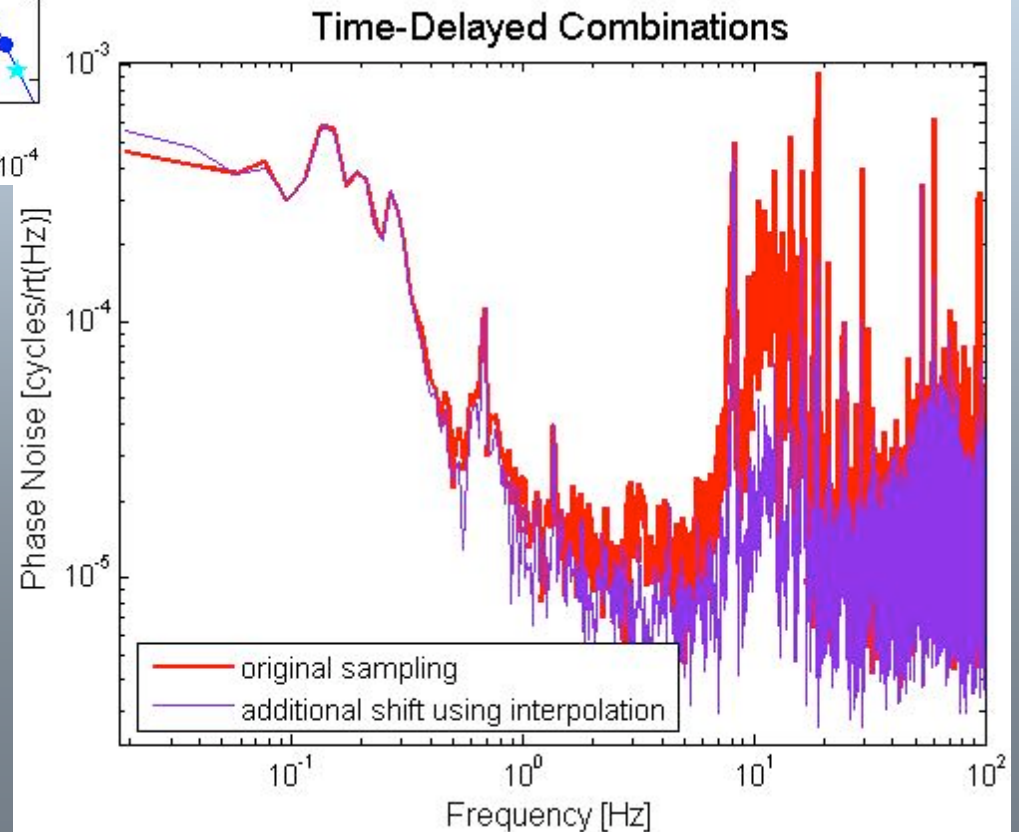


Interpolation

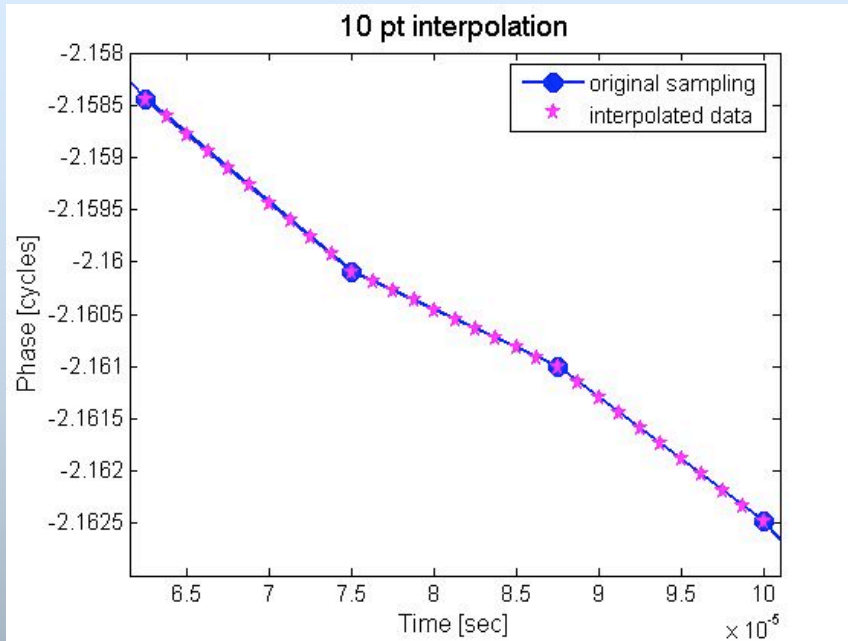
Used linear interpolation to
shift one data set 2 _sec



$$2.0005625 + 0.000002 = 2.0005645 \text{ sec}$$

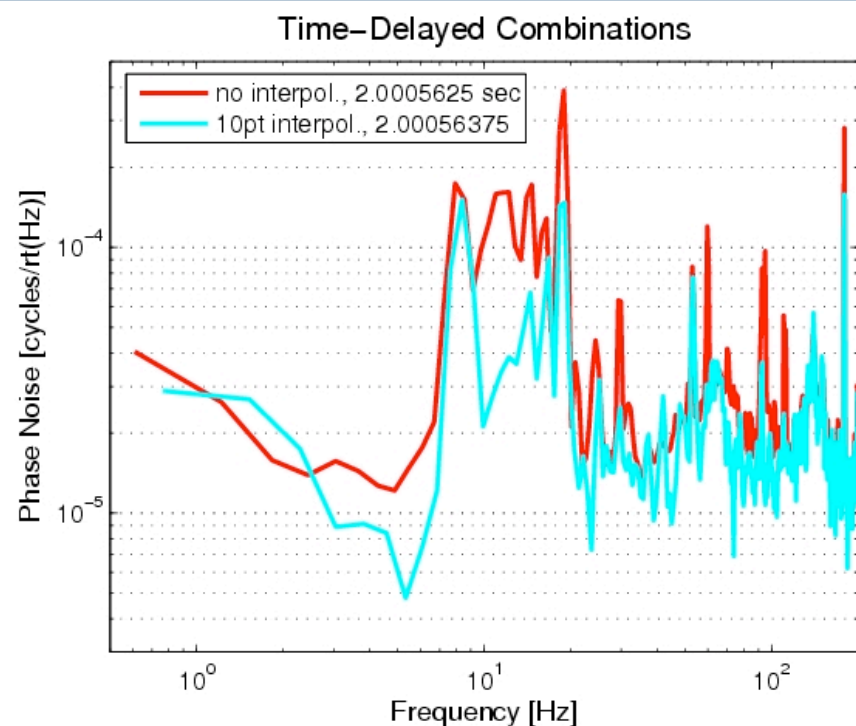


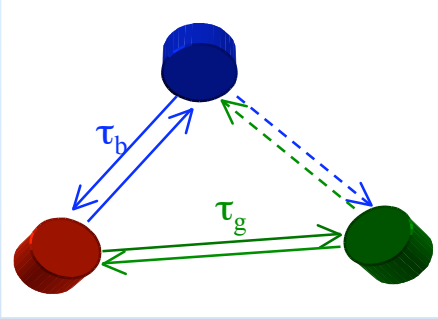
10-pt Interpolation



10-pt interpolation gives
1.25 μ sec resolution

The optimal shift of
2.00056375 sec is the
closest value to the
previous estimate of
2.00056425 sec.





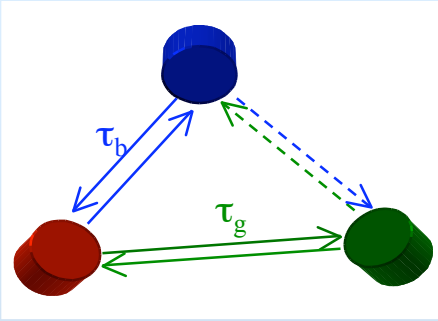
$$S_b(t) = -2\omega_0\tau_b - \phi(t) + \phi(t - 2\tau_b) + h_b(t) + h_b(t - \tau_b)$$

$$S_g(t) = -2\omega_0\tau_g - \phi(t) + \phi(t - 2\tau_g) + h_g(t) + h_g(t - \tau_g)$$

Instead of combining $S_b(t) - S_g(t)$, use $S_b(t) - S_g(t) - S_b(t - 2\tau_g) + S_g(t - 2\tau_b)$

$S_b(t)$:	$-2\tau_{0b}$	$-\phi(t)$	$+\phi(t - 2\tau_b)$		$h_b(t) + h_b(t - \tau_b)$
$-S_g(t)$:	$2\tau_{0g}$	$-\phi(t)$	$-\phi(t - 2\tau_g)$		$-h_g(t) - h_g(t - \tau_g)$
$-S_b(t - 2\tau_g)$:	$2\tau_{0b}$		$+\phi(t - 2\tau_g)$	$-\phi(t - 2\tau_b - 2\tau_g)$	$-h_b(t - 2\tau_g) - h_b(t - \tau_b - 2\tau_g)$
$S_g(t - 2\tau_b)$:	$-2\tau_{0g}$		$-\phi(t - 2\tau_b)$	$+\phi(t - 2\tau_b - 2\tau_g)$	$h_g(t - 2\tau_b) + h_g(t - \tau_g - 2\tau_b)$

Second Generation TDI compensates for effects due to space craft motion with additional terms.



$$S_b(t) = -2\omega_0\tau_b - \phi(t) + \phi(t - 2\tau_b) + h_b(t) + h_b(t - \tau_b)$$

$$S_g(t) = -2\omega_0\tau_g - \phi(t) + \phi(t - 2\tau_g) + h_g(t) + h_g(t - \tau_g)$$

Instead of combining $S_b(t) - S_g(t)$, use $S_b(t) - S_g(t) - S_b(t - 2\tau_g) + S_g(t - 2\tau_b)$

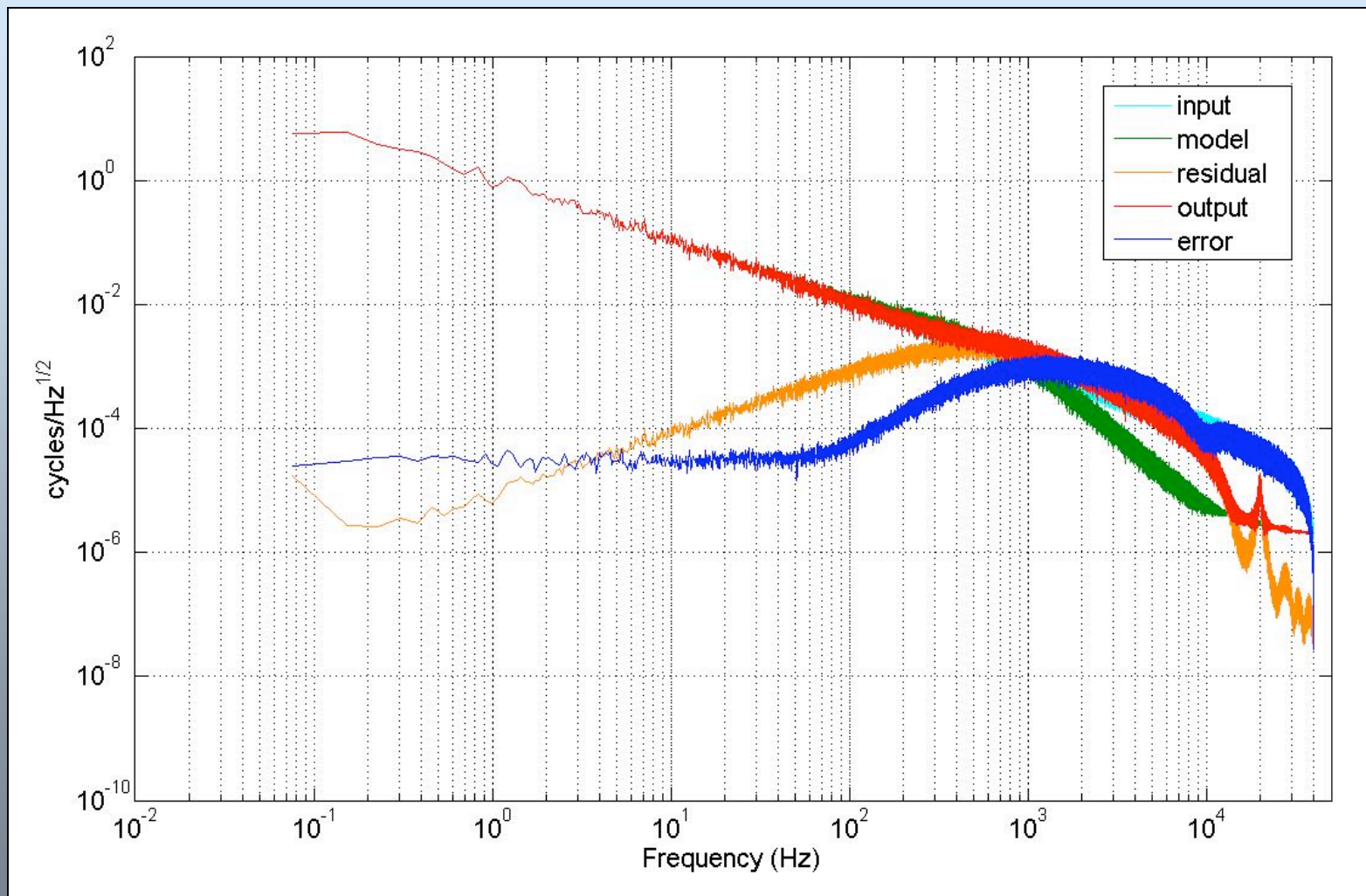
$S_b(t):$	$-2\tau_{0b}$	$-\phi(t)$	$\phi(t - 2\tau_b)$		$h_b(t) + h_b(t - \tau_b)$
$-S_g(t):$	$2\tau_{0g}$	$-\phi(t)$	$-\phi(t - 2\tau_g)$		$-h_g(t) - h_g(t - \tau_g)$
$-S_b(t - 2\tau_g):$	$2\tau_{0b}$		$\phi(t - 2\tau_g)$	$-\phi(t - 2\tau_b - 2\tau_g)$	$-h_b(t - 2\tau_g) - h_b(t - \tau_b - 2\tau_g)$
$S_g(t - 2\tau_b):$	$-2\tau_{0g}$		$-\phi(t - 2\tau_b)$	$\phi(t - 2\tau_b - 2\tau_g)$	$h_g(t - 2\tau_b) + h_g(t - \tau_g - 2\tau_b)$

All terms cancel except for those due to the change in phase from a gravitational wave!

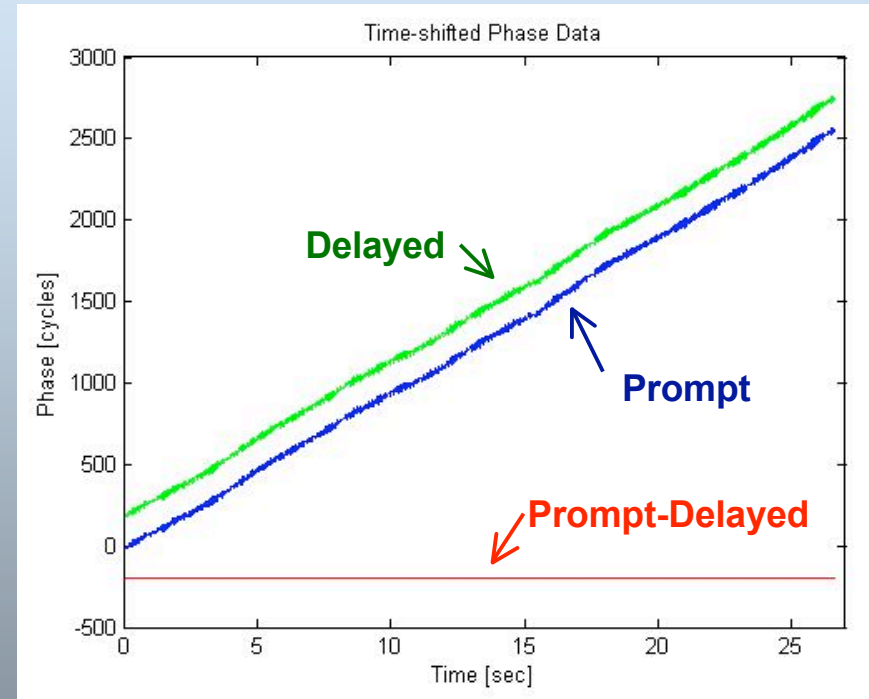
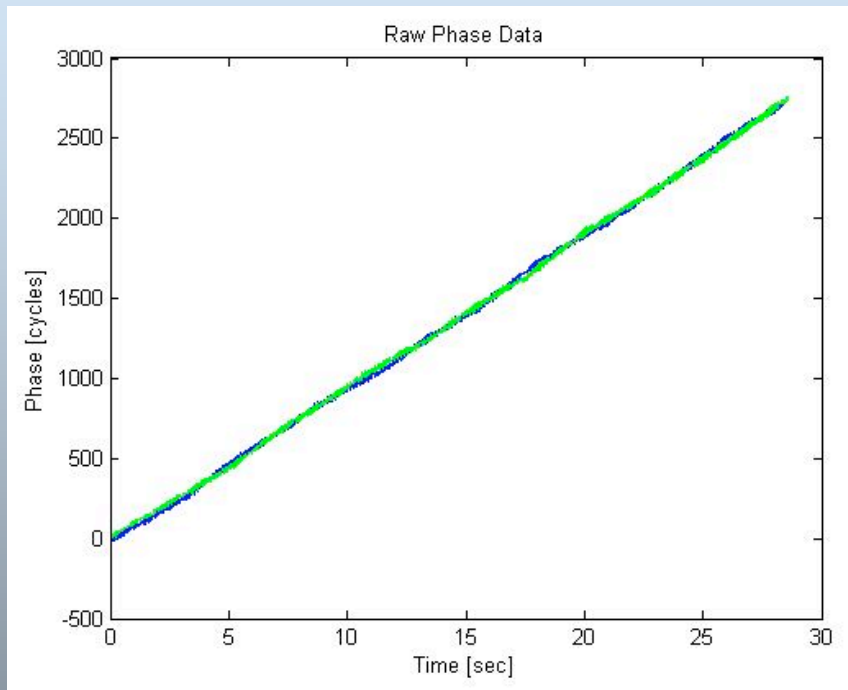


Second Generation TDI compensates for effects due to space craft motion with additional terms.

Modeled Performance of Post-Process Phasemeter



Initial Optical Experiment



Time-Delayed Comb. & Timing Res. Sup. Limit for Two-Arm Exp.

